

Black holes and fuzzballs

Oleg Lunin

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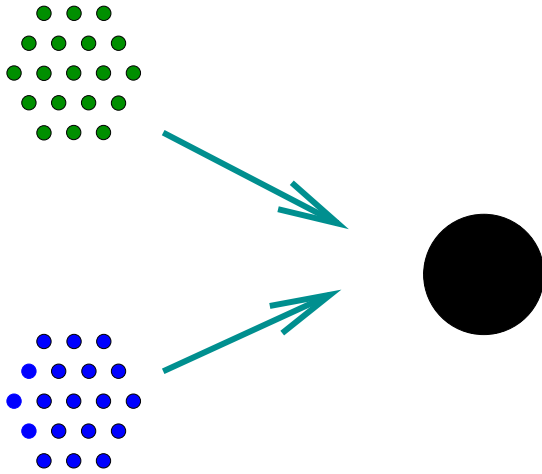
OL–Mathur–Turton, 1208.1770

Giusto–OL–Mathur–Turton, 1211.0306

OL–Mathur, 1211.5830

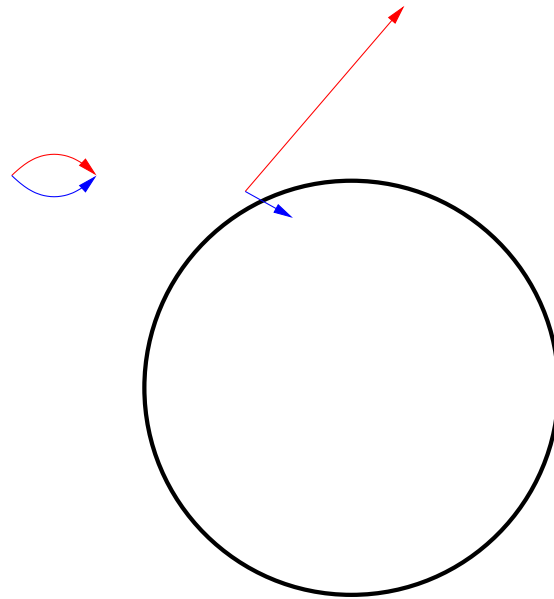
Black hole information paradox

- Black holes and Hawking radiation
 - classical theory: no-hair theorem



Black hole information paradox

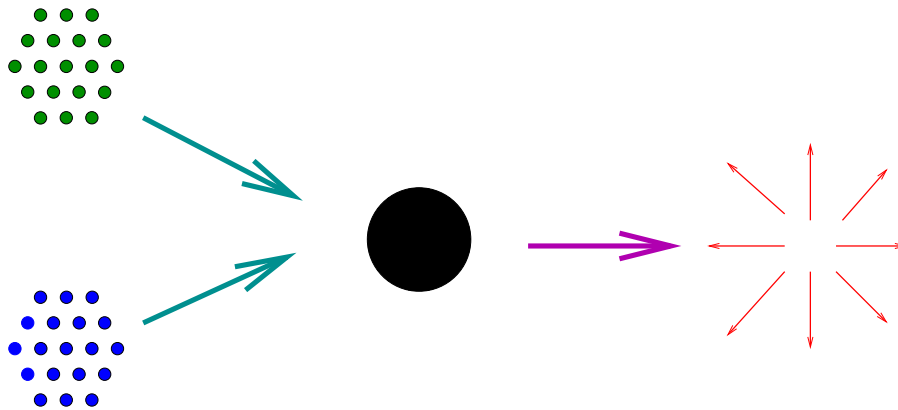
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 - QFT: pair creation



Hawking '74

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 - large number of microstates: $S = \frac{A}{4}$
 - paradox: pure state \Rightarrow mixed state

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- Options for resolving the paradox
 - information is lost
 - loophole in Hawking's argument

Outline

- Information paradox and AdS/CFT
 - evolution is unitary
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 - D1–D5–P states localized at the “horizon”
 - momentum modes on the torus
 - energy spectrum and CFT

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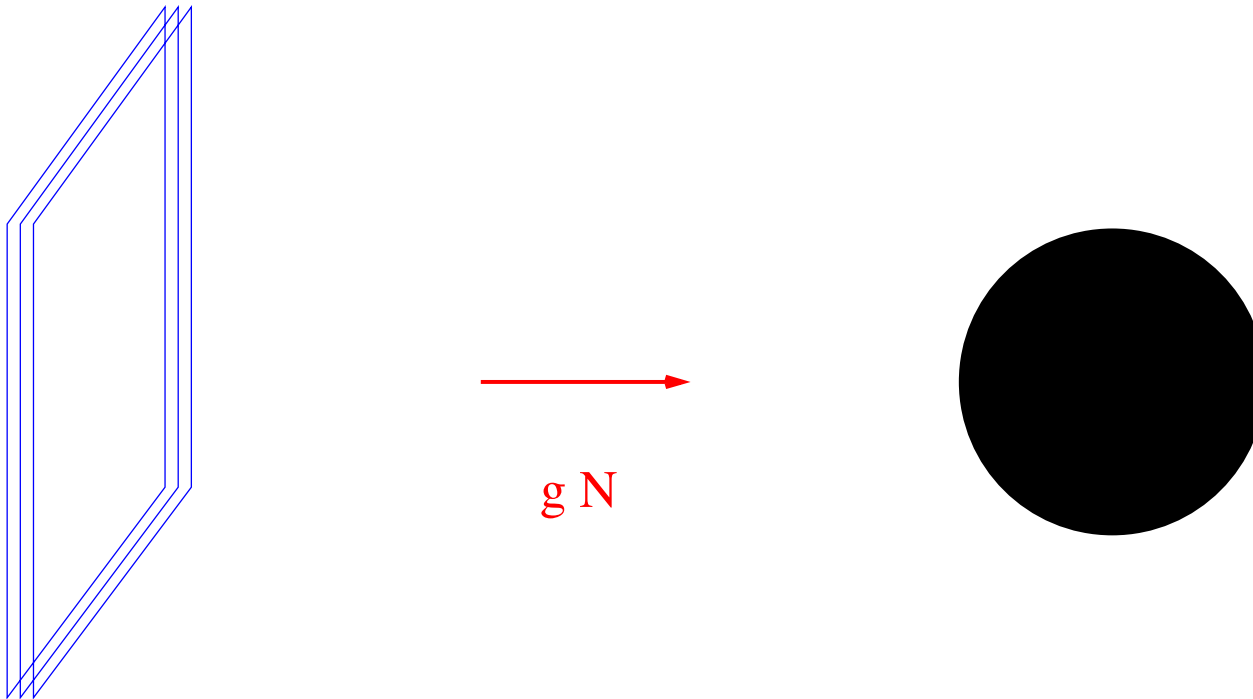
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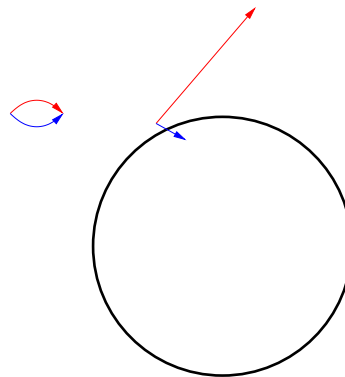
Maldacena '97

Black holes in AdS/CFT

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- AdS/CFT correspondence
 - unitary QFT \Rightarrow option B
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- Geometry must change **at the horizon**



Entropy and microscopic states

Entropy and microscopic states

- D1–D5–P system

	1	2	3	4	5	6	7	8	9
D1					●	~	~	~	~
D5					●	●	●	●	●
P					●	~	~	~	~

Entropy and microscopic states

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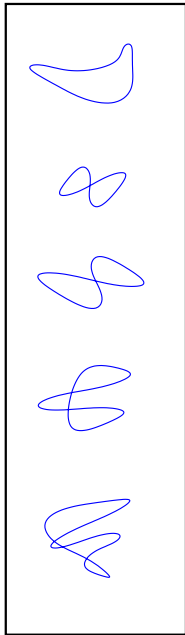
- Number of states at weak coupling:

$$S = k \ln \Gamma = \frac{A}{4\hbar G_N}$$

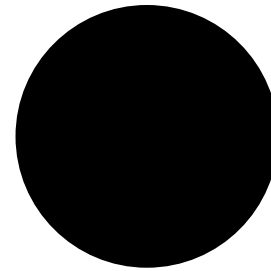
Strominger–Vafa '96

Entropy and microscopic states

● State counting



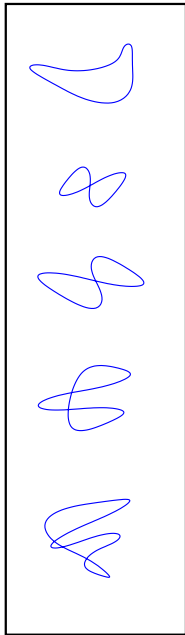
→
g N



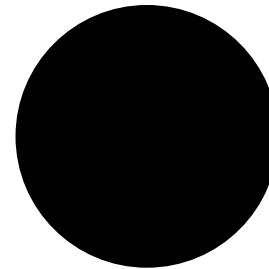
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g_N

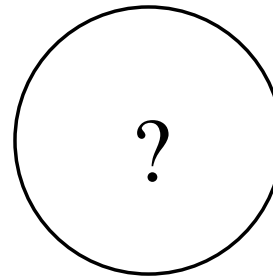


Strominger–Vafa '96

● Individual states

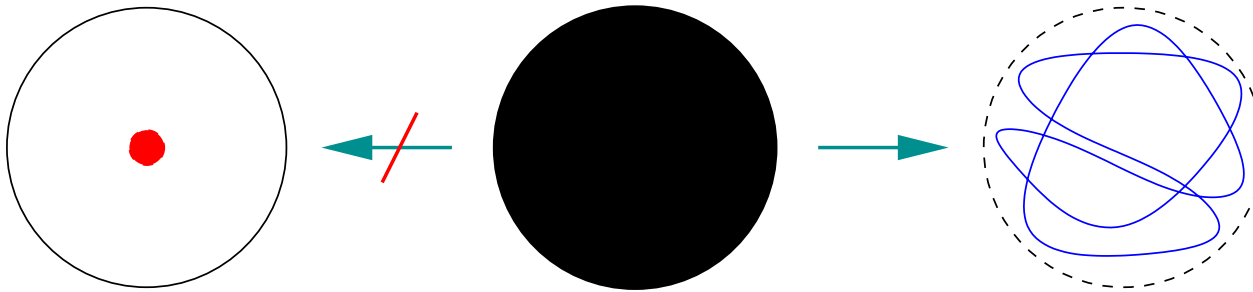


g_N



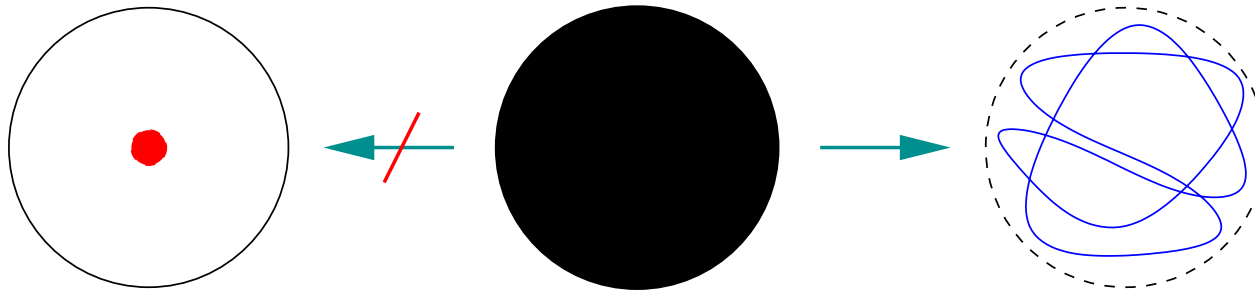
Fuzzballs

● Basic idea



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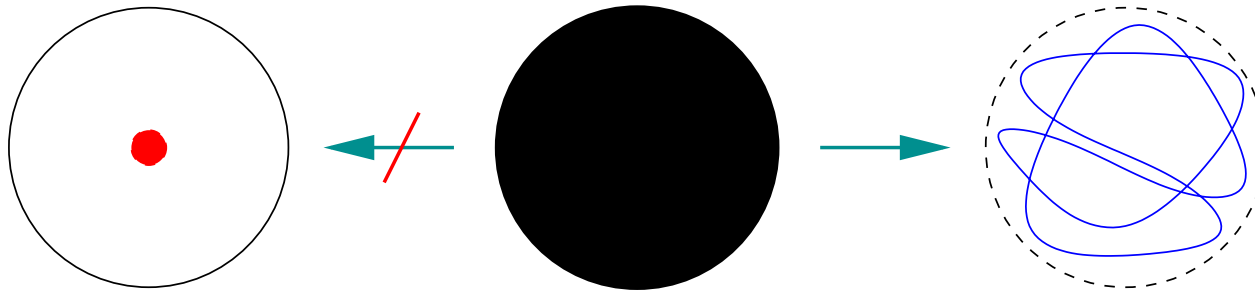


- individual horizon-free metric for every state
- distinctive features extending to the horizon
- corrections to Hawking radiation

Lunin–Mathur '01-02; Mathur '04

Fuzzballs

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Correspondence

Lunin–Mathur '01-02; Mathur '04

semiclassical state \Leftrightarrow

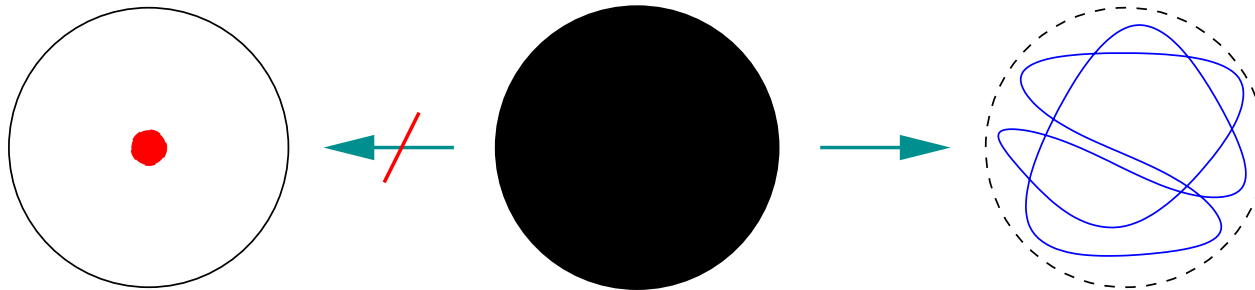
solution to SUGRA

generic state \Leftrightarrow

stringy geometry

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semiclassical state \Leftrightarrow solution to SUGRA
generic state \Leftrightarrow stringy geometry

● Horizon – coarse graining over geometries

Fuzzballs for D1–D5 system

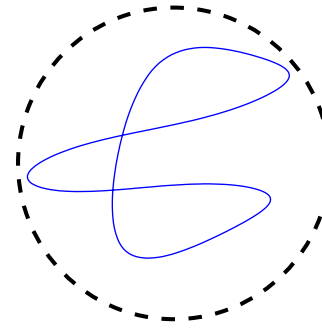
- Microstates and regular geometries

Fuzzballs for D1–D5 system

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$g N$

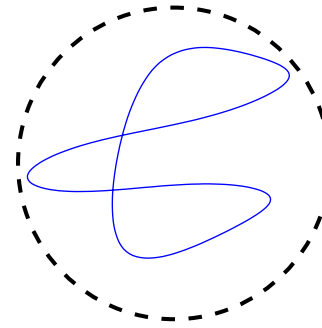


Fuzzballs for D1–D5 system

● Microstates and regular geometries



$g N$



vibrating string: $\mathbf{F}(v)$ \Leftrightarrow

geometry with

$$H_5 = 1 + \int \frac{dv}{(\mathbf{x} - \mathbf{F})^2}, \dots$$

OL–Mathur '01

vibration in 10D \Leftrightarrow

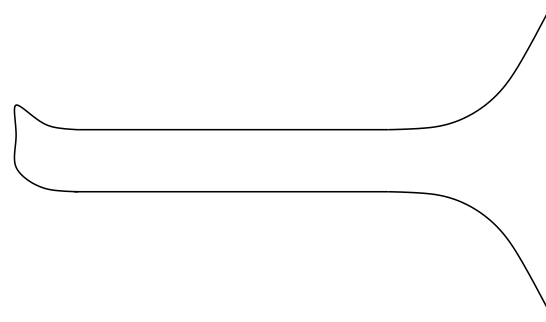
regular solution of SUGRA

OL–Maldacena–Maoz '02

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- Microstates and regular geometries

vibrating string $F(v)$ \Leftrightarrow regular solution of SUGRA



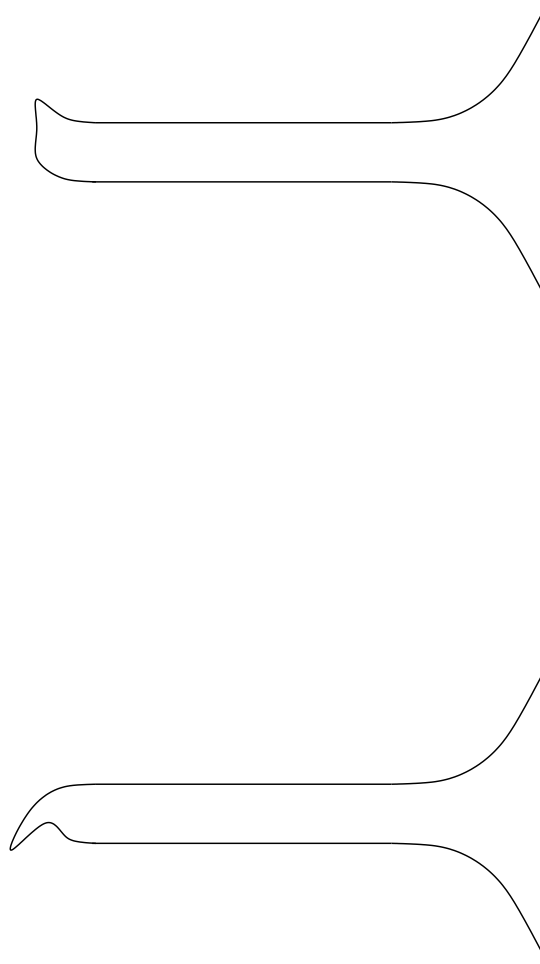
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r, S^3

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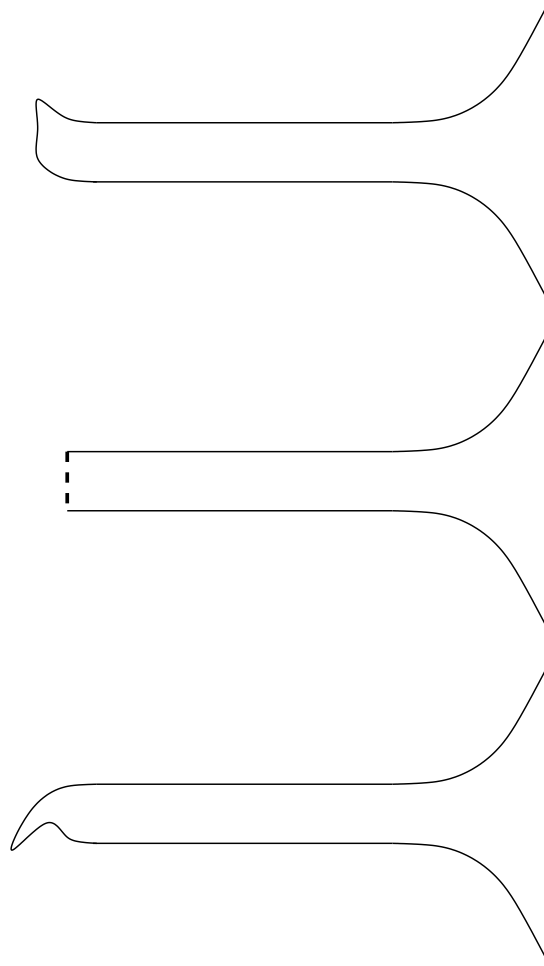
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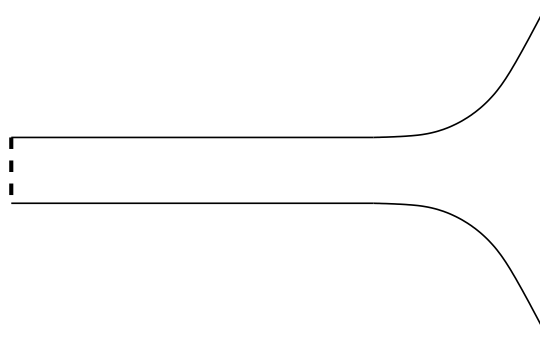
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Fuzzballs for D1–D5 system

- Microstates and regular geometries

vibrating string $F(v)$ \Leftrightarrow regular solution of SUGRA



- Horizon as coarse graining

- D1–D5: stretched horizon with Planckian area

Sen '95

- size of the fuzzball \sim size of the horizon

Lunin–Mathur '02

D1–D5–P and spectral flow

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 - geometries for all microstates
 - Planckian horizon

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$$T(z), \quad J^a(z), \quad G^{1,2}(z), \quad \bar{G}_{1,2}(z)$$

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$$T(z), \quad J^a(z), \quad G^{1,2}(z), \quad \bar{G}_{1,2}(z)$$

- boundary conditions for fermions

$$G(z) = -e^{i\pi\eta} G(e^{2\pi i} z), \quad J^\pm(z) = e^{\mp 2\pi i \eta} J^\pm(e^{2\pi i} z)$$

Schwimmer–Seiberg '87

D1–D5–P and spectral flow

- D1–D5 system
 - geometries for all microstates
 - Planckian horizon
- Momentum and spectral flow in $N = (4, 4)$ SCFT₂
 - spectral flow in CFT \Leftrightarrow diff in AdS
 - mixture between AdS_3 and S^3 ($\psi \rightarrow \psi + \eta(t - y), \dots$)

$$ds^2 = L^2 \left[-(r^2 + 1)dt^2 + \frac{dr^2}{r^2 + 1} + r^2 dy^2 + d\theta^2 + \cos^2 \theta d\psi^2 + \sin^2 \theta d\phi^2 \right]$$

- momentum charge is generated
- Integer spectral flow is a regular diffeomorphism

D1–D5–P and spectral flow

- Momentum and spectral flow in $N = (4, 4)$ SCFT₂
 - spectral flow in CFT \Leftrightarrow diff in AdS
 - momentum charge is generated
- Adding momentum to D1–D5 system
 - remove flat asymptotics
 - perform spectral flow
 - attach flat asymptotics

OL '04; Giusto–Mathur–Saxena '04

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- More states of D1–D5–P system
 - general solution in terms of 4D hyper–Kähler base Gutowski–Martelli–Reall '03
 - special **regular** solutions Bena–Warner–Bobev–Wang–... '05-11

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 - microscopic interpretation?

Fractional spectral flow

- CFT dual of the D1–D5 system
 - marginal deformation of the free CFT on orbifold
 - orbifold point: $(T^4)^N / S_N$, $N = n_1 n_5$

de Boer '98; Seiberg–Witten '99;

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 - chiral primaries: twist operators $\sigma_k(z)$

$$X_1 \rightarrow X_2 \rightarrow \dots X_k \rightarrow X_1$$

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- functions are single-valued on the covering space
- microstate with $\sigma_k \Rightarrow$ SF on the cover \Rightarrow FSF

$$\psi(z) = -e^{i\pi\eta} \psi(e^{2\pi i k} z)$$

- gravity: $[\sigma_k]^{N/k} \Rightarrow$ conical singularity

Fractional spectral flow

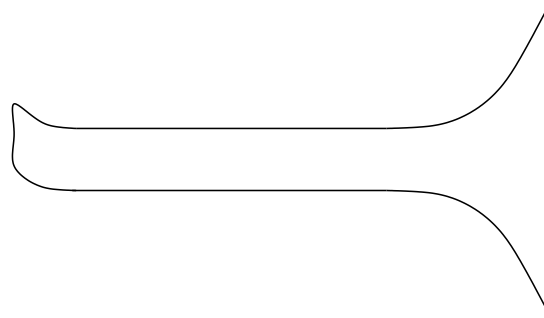
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- Fractional spectral flow in gravity
 - AdS region:

$$\psi \rightarrow \psi + \frac{s}{k}(t - y), \quad \phi \rightarrow \phi - \frac{s}{k}(t - y)$$

- flat asymptotics – GMR or limit of a black hole
- conical singularity associated with σ_k softens

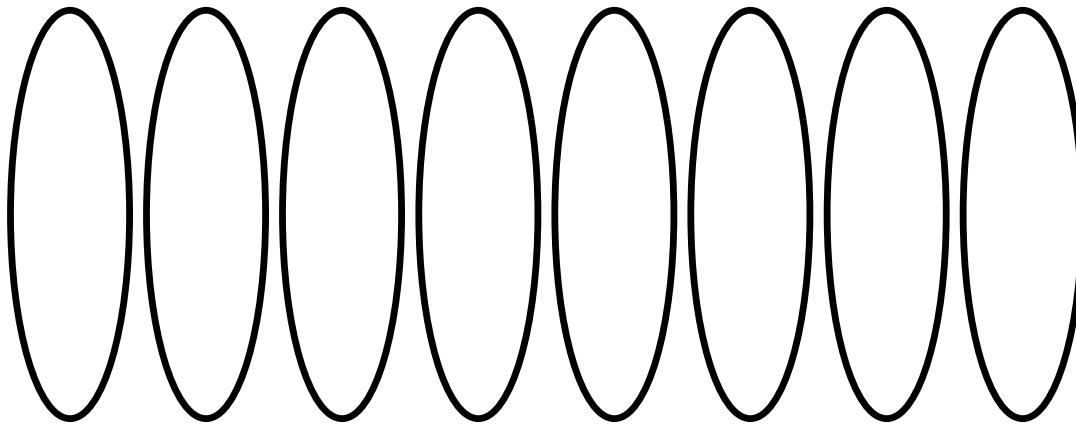
Properties of “fractional microstates”

- Geometry is modified in the “cap”



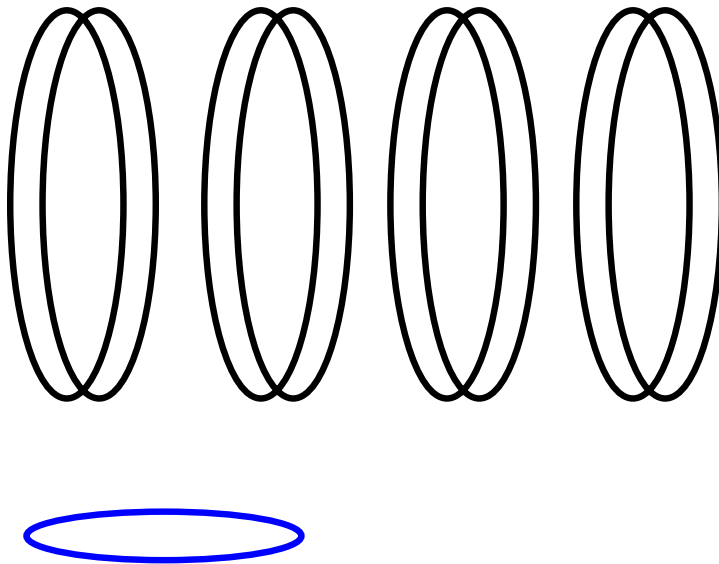
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 - CFT picture — multiwound string



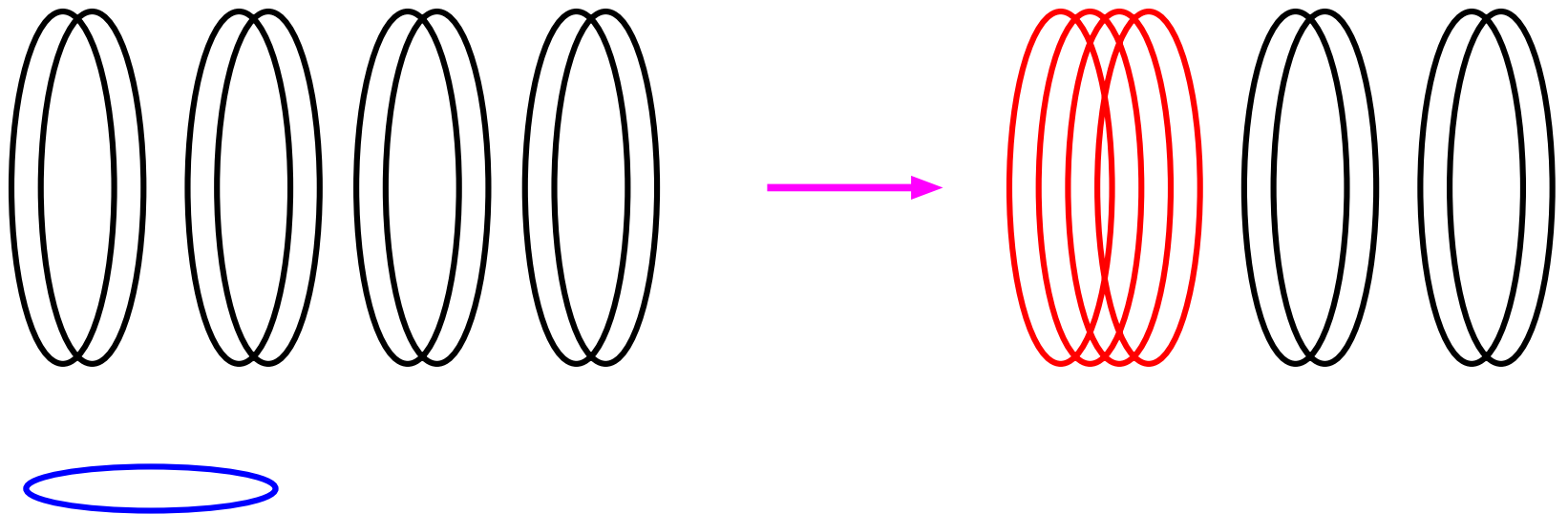
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- Contributions to the energy
 - fermions on a longer string $\Rightarrow \Delta h_f = -\frac{ls}{k}$
 - bosons from the vertex operator $\Rightarrow \Delta h_b = \frac{1}{k}$
 - ang. mom. from the vertex operator: $\Delta h_J = \frac{m(2s+1)}{k}$
 - momentum of the excitation: $n_p = \bar{h} - h$

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- Agreement with SUGRA

Giusto–OL–Mathur–Turton '12

$$\omega = \frac{l + 2 + 2(2s + 1)m}{k} - n_p$$

Excitations on the torus

- Adding momentum to D1–D5
 - spectral flow involves $S^3 \Rightarrow$ solution of 6D SUGRA
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 - start with known microstate
 - solve for linear perturbation
 - guess the solution
 - check the eqns for Killing spinors

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- Deformation of GMR: OL–Mathur–Turton '12

$$\begin{aligned} ds^2 &= -e^{\hat{v}} e^{\hat{u}} + H ds_4^2 + dz^\alpha dz^\alpha, \\ C^{(2)} &= \frac{1}{2} e^{\hat{v}} \wedge e^{\hat{u}} + e^{\hat{v}} \wedge \mathcal{A} + \sigma_2, \\ e^{\hat{v}} &= H^{-1}(dv + \beta), \quad e^{\hat{u}} = du + \omega + \frac{FH}{2} e^{\hat{v}} + \Phi_\alpha(v, x) dz^\alpha. \end{aligned}$$

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 - linear equation for the deformation
$$d \left(*_6 \left[D\Phi_\alpha + \partial_v \Phi_\alpha H e^{\hat{v}} \right] \right) = 0$$
 - regular GMR geometry \Rightarrow regular deformation
 - CFT excitation on the boundary \Rightarrow unique solution

Relation to other constructions

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- Plane wave on static black hole

- Garfinkle–Vachaspati transform: $g_{\mu\nu} \rightarrow g_{\mu\nu} + \Phi k_\mu k_\nu$
- application to black holes \Rightarrow singularity

Horowitz–Marolf '96

- GV is extended to some non–static geometries

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 - GV is extended to some non–static geometries
- Pp–wave on D1–D5–P via dualities

$$ds^2 = -e_0^{\hat{v}} e_0^{\hat{u}} - \frac{1}{4} \Phi_\alpha \Phi_\alpha (e_0^{\hat{v}})^2 + H ds_4^2 + dz^\alpha dz^\alpha,$$

$$B = -\frac{1}{2} \Phi_\alpha e_0^{\hat{v}} \wedge dz^\alpha, \quad C^{(4)} = \frac{1}{24} \left[e_0^{\hat{v}} \wedge \varepsilon_{\alpha\beta\gamma\delta} \Phi_\alpha dz^\beta \wedge dz^\gamma \wedge dz^\delta + dual \right],$$

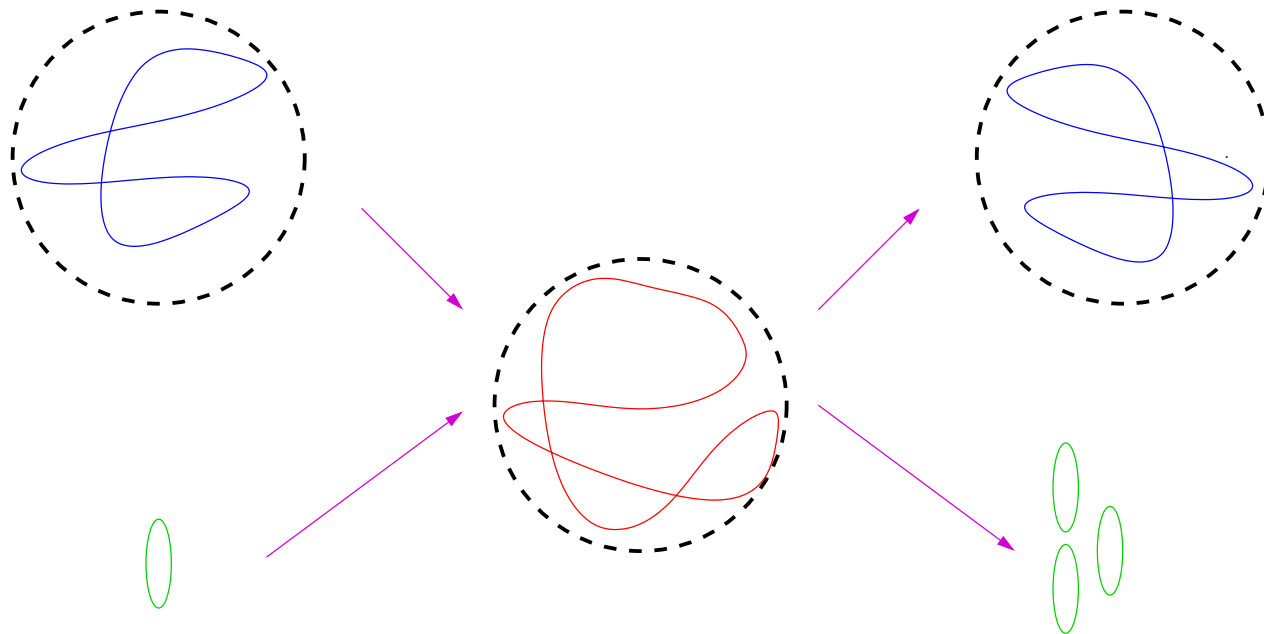
$$C^{(6)} = \left[\frac{1}{2} e_0^{\hat{v}} \wedge e_0^{\hat{u}} + e_0^{\hat{v}} \wedge \mathcal{A} + \sigma_2 \right] dz_1 \wedge dz_2 \wedge dz_3 \wedge dz_4.$$

Hawking radiation from CFT

- Non-extremal BH: radiation in probe approximation

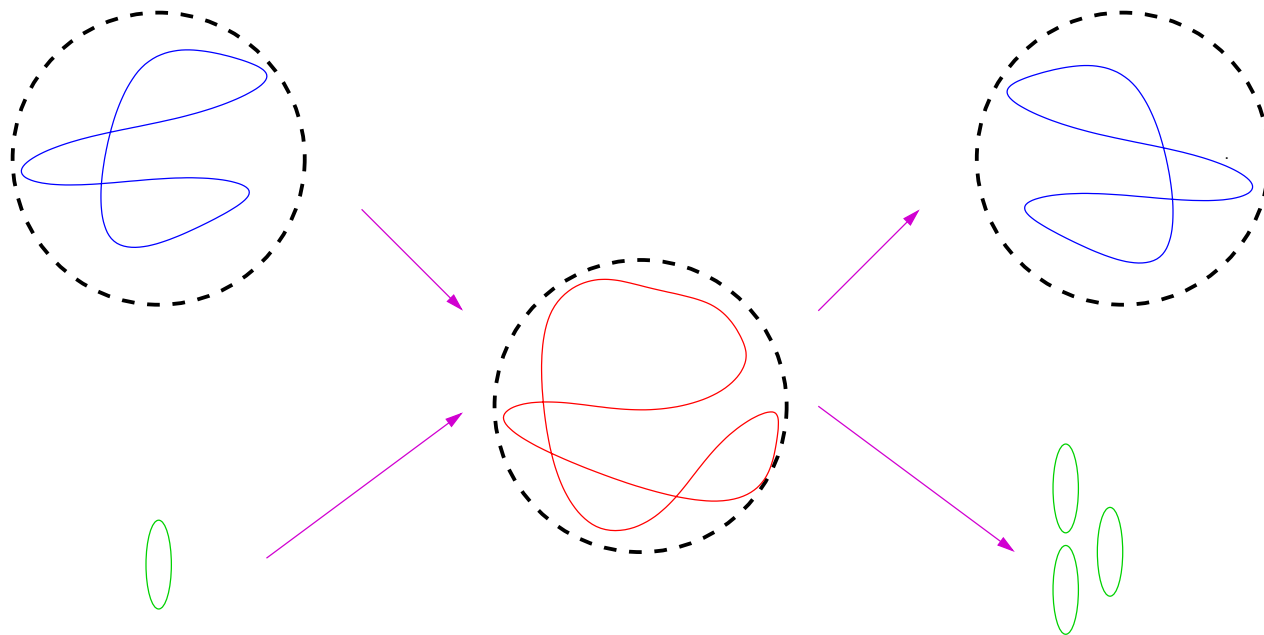
Hawking radiation from CFT

- Non-extremal BH: radiation in probe approximation
- Hawking from near-extremal fuzzballs
 - excitation of the microstate
 - transition to a different microstate via emission



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 - orbifold CFT with spin operators (R-sector)

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Lunin–Mathur '12

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 - one outgoing particle
 - orbifold CFT with spin operators (R–sector)
- Results for the S–matrix
 - explicit scattering amplitudes
 - transitions to “nearby” states are preferred
- Open questions
 - several outgoing particles/general microstates
 - gravity counterpart of the calculation

Lunin–Mathur '12

Summary

- Fuzzball proposal for black holes
 - individual microstate \Leftrightarrow regular (stringy) geometry
 - interior of the "horizon" is not empty
 - horizon as a result of the coarse graining
 - complete picture for D1–D5 system

Summary

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 - generic microstates, Hawking radiation in gravity...